

REPORT OF REFEREE

Stefi A. Baum et al.: EXTENDED OPTICAL LINE EMITTING GAS IN  
Author, Title ~~RADIO GALAXIES: BROAD-BAND OPTICAL, NARROW-BAND OPTICAL, AND~~  
RADIO IMAGING OF A REPRESENTATIVE SAMPLE

This is an impressive piece of work and the authors are to be commended for publishing it in the Supplements instead of piecemeal.

I have noted a few places in the manuscript where a change or addition is needed.

It seems important for the authors to refer to Hansen, Nørgaard-Nielson, and Jørgensen (December) 1987, A&A Supplements article, which covers much of the same optical material.

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TABLE 1  
 Sample Members

| Source        | Redshift           | RA          | DEC                      | $E_{B-V}$        | kpc arcsec <sup>-1</sup> |
|---------------|--------------------|-------------|--------------------------|------------------|--------------------------|
| 3C 29*        | 0.0447             | 00 55 01.56 | -01 39 39.9              | 0.043            | 0.81                     |
| 3C 33*        | 0.0595             | 01 06 14.55 | +13 04 15.4              | 0.03             | 1.06                     |
| 3C 40*        | 0.0177             | 01 23 27.55 | -01 36 18.9 <sup>b</sup> | 0.043            | 0.33                     |
| B2 0149 + 358 | 0.0162             | 01 49 50.01 | +35 54 20.5 <sup>b</sup> | 0.04             | 0.31                     |
| 3C 63*        | 0.175 <sup>c</sup> | 02 18 21.94 | -02 10 32.5              | 0.0              | 2.68                     |
| 3C 75N*       | 0.0241             | 02 55 02.99 | +05 49 37.0 <sup>b</sup> | 0.09             | 0.45                     |
| 3C 75S*       | 0.0241             | 02 55 03.08 | +05 49 20.9 <sup>b</sup> | 0.09             | 0.45                     |
| 3C 78*        | 0.0288             | 03 05 49.05 | +03 55 13.1 <sup>b</sup> | 0.09             | 0.54                     |
| 3C 88*        | 0.0302             | 03 25 18.20 | +02 23 20.0              | 0.09             | 0.56                     |
| 3C 89*        | 0.1386             | 03 31 43.01 | -01 20 55.9              | 0.06             | 2.22                     |
| PKS 0349-278* | 0.066              | 03 49 31.77 | -27 53 31.3              | ...              | 1.16                     |
| 3C 98*        | 0.0306             | 03 56 10.17 | +10 17 32.5              | 0.15             | 0.57                     |
| 3C 105*       | 0.0886             | 04 04 38.99 | +03 34 27.5              | 0.15             | 1.52                     |
| 3C 109*       | 0.3056             | 04 10 54.86 | +11 04 40.9              | 0.21             | 4.02                     |
| PKS 0634-206* | 0.056              | 06 34 23.27 | -20 32 18.5              | 0.50             | 1.00                     |
| 3C 171        | 0.2384             | 06 51 10.92 | +54 12 48.1 <sup>b</sup> | 0.09             | 3.38                     |
| PKS 0745-191* | 0.1028             | 07 45 18.45 | -19 10 11.6 <sup>a</sup> | 0.3 <sup>a</sup> | 1.73                     |
| 3C 192*       | 0.0598             | 08 02 32.31 | +24 18 54.9 <sup>b</sup> | 0.03             | 1.06                     |
| 3C 196.1*     | 0.198              | 08 12 56.99 | -02 59 13.7 <sup>a</sup> | 0.06             | 2.94                     |
| 3C 218*       | 0.055              | 09 15 41.20 | -11 53 04.9              | 0.03             | 0.98                     |
| 3C 219*       | 0.1744             | 09 17 50.68 | +45 51 43.8              | 0.0              | 2.67                     |
| 3C 223*       | 0.1368             | 09 36 50.86 | +36 07 35.5 <sup>a</sup> | 0.0              | 2.20                     |
| 3C 227*       | 0.0861             | 09 45 06.54 | +07 39 17.4              | ...              | 1.48                     |
| 3C 264*       | 0.0208             | 11 42 29.56 | +19 53 02.3              | 0.0              | 0.39                     |
| 3C 272.1*     | 0.0031             | 12 22 31.58 | +13 09 50.7              | 0.02             | 0.06                     |
| 3C 274*       | 0.0043             | 12 28 17.56 | +12 40 02.0              | 0.02             | 0.07                     |
| 3C 275*       | 0.480              | 12 39 44.91 | -04 29 53.9 <sup>a</sup> | 0.007            | 5.28                     |
| 3C 278*       | 0.0145             | 12 51 58.56 | -12 17 51.9              | ...              | 0.28                     |
| 3C 285        | 0.0704             | 13 10 05.99 | +14 50 55.7 <sup>b</sup> | ...              | 1.99                     |

axis.

- 3. The presence of much fainter line emitting gas  $\sim 45''$  (25 kpc) northeast of the nucleus, along the radio source axis, and of a very faint, curving, tangential filament of line emitting gas  $\sim 30''$  (17 kpc) due south of the nucleus, both of which are confirmed spectroscopically.

*Note  
discussion  
of excitation  
conditions  
in SIMKIN 1979*

In the radio, at  $3.8''$  resolution, 3C 98 shows single hotspots in both the northern and the southern radio lobes, separated by  $\sim 310''$  (175 kpc). A jet is present in the northern lobe, though it is strongly blended with diffuse lobe emission in the image shown in Figure 14(b). The jet is straight along most of its length, but curves abruptly to the east at a bright knot  $\sim 120''$  from the core, before merging into the northern hotspot. There is a ridge of bright emission curving to the south of the northern hotspot. The large scale radio structure of 3C 98 is under represented in our maps, but 3C 98 does show a cocoon or bridge of lower surface brightness radio emission extending from the hotspots back towards the core which, at least to the north, appears to show a tail of emission extending to the west of the radio source axis, near the radio core. The fractional polarization of the radio emission from 3C 98 at  $3.8''$  resolution and 6 cm is high throughout most of the source structure. The average fractional polarization over the source is 28%.

An understanding of the nature of the relationship between the extended line emitting gas and the radio source structure awaits more detailed studies. As described above, the brightest extra-nuclear line emitting gas is skewed with respect to the radio source axis but the most distant line emitting gas is located along the radio source axis.

### 3C 105 (0404+03)

The host galaxy of 3C 105 is a surprisingly small and underluminous galaxy to be associated with a high luminosity radio source. We detect only nuclear  $H\alpha + [NII]$  emission from the host galaxy of 3C 105. In the radio, 3C 105 shows a bright southern hotspot and only a weak, resolved hotspot in the northern lobe (at  $6''$  resolution), separated by  $\sim 335''$  (510 kpc). Narrow bridges of emission extend from the hotspots back towards the core.

### 3C 109 (04120+11)

The host galaxy of 3C 109 is an isolated, broad line (Grandi and Osterbrock 1978), N